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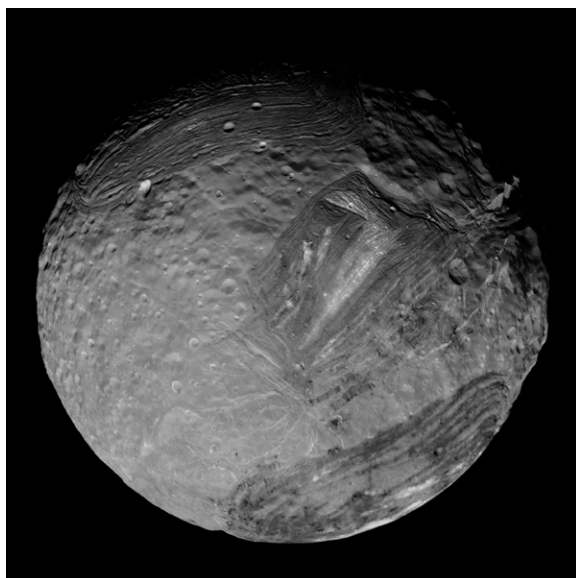
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COVER: Cover: Mosaic of Voyager 2 images of Miranda, a small but intensely tectonically deformed icy satellite of Uranus. Three zones of endogenic resurfacing, called coronae, may have formed as a result of convective upwelling driving concentric surface extension. See "Global resurfacing of Uranus's moon Miranda by convection" by Hammond and Barr, p. 931–934.

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